# BALANCED TRUNCATION OF LINEAR SECOND-ORDER SYSTEMS: A HAMILTONIAN APPROACH 

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#### Abstract

We present a formal procedure for structure-preserving model reduction of linear second-order and Hamiltonian control problems that appear in a variety of physical contexts, e.g., vibromechanical systems or electrical circuit design. Typical balanced truncation methods that project onto the subspace of the largest Hankel singular values fail to preserve the problem's physical structure and may suffer from lack of stability. In this paper we adopt the framework of generalized Hamiltonian systems that covers the class of relevant problems and that allows for a generalization of balanced truncation to second-order problems. It turns out that the Hamiltonian structure, stability and passivity are preserved if the truncation is done by imposing a holonomic constraint on the system rather than standard Galerkin projection.


Key words. structure-preserving model reduction, generalized Hamiltonian systems, balanced truncation, strong confinement, invariant manifolds, Hankel norm approximation.

AMS subject classifications. 70Q05, 70J50, 93B11, 93C05, 93C70.

1. Introduction. Model reduction is a major issue for control, optimization and simulation of large-scale systems. In particular for linear time-invariant systems, balanced truncation is a well-established tool for deriving reduced (i.e., low-dimensional) models that have an input-output behaviour similar to the original model [1, 2]; see also [3] and the referenced therein. The general idea of balanced truncation is to restrict the system onto the subspace of easily controllable and observable states which can be determined by the computing the Hankel singular values associated with the system. Moreover the method is known to preserve certain properties of the original system such as stability or passivity and gives an error bound that is easily computable. One drawback of balanced truncation is that there is no straightforward generalization to second-order systems; see, e.g., [4] or the recent articles [5, 6, 7, 8] for a discussion of various possible strategies. Second-order equations occur in modelling and control of many physical systems, e.g., electrical circuits, structural mechanics or vibro-acoustic models; see, e.g., $[9,10]$, and the main objective of extending balancing methods to such systems is to derive reduced models that remain physically interpretable, i.e., that inherit the interaction structure of the original model. Ordinary balanced truncation proceeds by recasting the equations in first-order form, and we shall argue that an obvious (and in some sense natural) first-order formulation of a second-order control system that is amenable to balancing takes into account the system's Hamiltonian structure. This leads to a formulation of the control problem as a generalized Hamiltonian system [11]. Nonetheless balanced truncation, if applied to Hamiltonian system, though preserving stability or passivity and giving nicely computable error bounds, fails to preserve the underlying Hamiltonian structure.
[^0]Main result. Given a quadratic Hamiltonian $H: \mathbf{R}^{2 n} \rightarrow \mathbf{R}$, we consider linear state space systems that are of the general form

$$
\begin{aligned}
\dot{x}(t) & =(J-D) \nabla H(x(t))+B u(t), \quad x(0)=x \\
y(t) & =C \nabla H(x(t))
\end{aligned}
$$

where $J \in \mathbf{R}^{2 n \times 2 n}$ is the invertible skew-symmetric structure matrix, $D \in \mathbf{R}^{2 n \times 2 n}$ is the symmetric positive semidefinite friction matrix, $B \in \mathbf{R}^{2 n \times m}$ and $C \in \mathbf{R}^{l \times 2 n}$ are the coefficients for control $u \in \mathbf{R}^{m}$ and output $y \in \mathbf{R}^{l}$. It turns out that balanced truncation can be extended to the aforementioned class of Hamiltonian system while preserving the Hamiltonian structure and properties such as stability or passivity. More precisely, letting $S \subset \mathbf{R}^{2 n}$ denote the $d$-dimensional subspace of the most controllable and observable states, we prove, using a perturbation argument, that we can replace the full system by the dimension-reduced system

$$
\begin{aligned}
& \dot{\xi}(t)=\left(\tilde{J}_{11}-\tilde{D}_{11}\right) \nabla \bar{H}(\xi(t))+\tilde{B}_{1} u(t), \quad \xi(0)=\xi \\
& y(t)=\tilde{C}_{1} \nabla \bar{H}(\xi(t))
\end{aligned}
$$

that yields a "good" approximation of the output variable $y$. The idea of our approach is to consider the limit of vanishing small Hankel singular values (associated with the negligible subspace $\mathbf{R}^{2 n} \backslash S$ ) which naturally leads to confinement of the dynamics to $S$. Accordingly the reduced coefficients $\tilde{J}_{11}, \tilde{D}_{11}, \tilde{B}_{1}, \tilde{C}_{1}$ are simply the former structure, friction, control and output matrices restricted to the subspace $S$, and $\bar{H}$ is an effective Hamiltonian that can be computed from the Schur complement of the matrix block $\tilde{E}_{11}$ that is the projection of $E=\nabla^{2} H$ onto $S$ (see Sec. 3.1).

We show that the reduced system is a again stable and passive state space system if the original system was and that the associated transfer function satisfies the usual balanced truncation error bound (see Sec. 5). It is important to note that our perturbation analysis is carried out not for the transfer function (i.e. in the frequency domain), but rather for the corresponding equations of motion (i.e., in the time domain). In other words, we derive approximations of the original dynamical system for any given initial condition whereas, e.g., standard balanced truncation yield only approximation for zero initial condition $x(0)=0$. To the best of our knowledge a systematic multiscale analysis of the equations of motion in the limit of vanishing Hankel singular values is new; see, e.g., $[12,13,14]$ for approaches in which low-rank perturbative approximations of the transfer function are sought.

Related strategies. We compare our confinement approach to two alternative, structure-preserving methods (Sec. 3.2 and the preceding paragraph) that have been proposed in [15] and the first of which uses a perturbation-like argument. Although both methods preserve stability and passivity it turns turns out that they do a bad job in terms of the balanced truncation error bound. In either case (including confinement) going back to the original second-order form is not possible in general as the reduced Hamiltonian is not a separable sum of potential and kinetic energy; a special situation in which the second-order structure may be preserved occurs when the balancing transformation decays into purely position and momentum (velocity) parts, and, in fact, alternative approaches such as [4, 6, 7] proceed by treating position and momentum (velocity) parts separately, but they also fail to preserve stability of the second-order system as has been pointed out in the recent work [7].

Preserving the Hamiltonian structure plus stability and passivity, however, has value in its own right as Hamiltonian models are prevalently used in e.g., multibody
dynamics [11] or circuit design [16]. Especially for the circuit systems, passivitypreserving reduction strategies based on (split) congruence transformations have been proposed in $[17,18]$ and, more recently, in [8]. Though no provable error bounds are available, numerical evidence suggests that all these methods give reasonable approximations of the corresponding transfer functions.

Outline of the article. In Section 2 we introduce the conceptual framework of generalized Hamiltonian systems and balancing. Section 3 deals with the problem of restricting the equations of motion to the most controllable and observable subspace by (holonomic) constraints; for this purpose we employ a singular perturbation argument and prove that the dynamics are confined to the essential subspace as the negligible Hankel singular values go to zero (Sec. 3.1); conditions under which the original second-order structure is preserved are discussed at the end of the section. Stability of the reduced systems is a subtle issue and we have devoted a separate section to it (Sec. 4). Last but not least, we prove in Section 5 that the transfer function associated with the original system converges in $H^{\infty}$ to the transfer function of the constrained system in the limit of vanishing small Hankel singular values, and we briefly discuss error bounds. The article concludes with two numerical examples in Section 6.
2. Set-up: generalized Hamiltonian systems. Given a smooth Hamiltonian

$$
H: \mathbf{R}^{2 n} \supseteq \mathbf{X} \rightarrow \mathbf{R}, \quad H(x)=\frac{1}{2} x^{T} E x
$$

with $E=E^{T} \succ 0$ (" $\succ$ " means positive definite) we consider systems of the form

$$
\begin{align*}
\dot{x}(t) & =(J-D) \nabla H(x(t))+B u(t) \\
y(t) & =C \nabla H(x(t)) \tag{2.1}
\end{align*}
$$

where $J \in \mathbf{R}^{2 n \times 2 n}$ is an invertible skew-symmetric matrix, $D \in \mathbf{R}^{2 n \times 2 n}$ is symmetric positive semidefinite, $B \in \mathbf{R}^{2 n \times m}$ and $C \in \mathbf{R}^{l \times 2 n}$ (all constant). We suppose that both the control function $u(\cdot) \in \mathbf{R}^{m}$ and the observable $y(\cdot) \in \mathbf{R}^{l}$ are in $L^{2}(\mathbf{R})$. For $C=B^{T}$ systems of type (2.1) are called port-Hamiltonian (see [11]).

Second-order systems. As can be readily checked, the second-order system

$$
\begin{align*}
M \ddot{q}(t)+R \dot{q}(t)+K q(t) & =B_{2} u(t) \\
y(t) & =C_{1} q(t)+C_{2} \dot{q}(t) \tag{2.2}
\end{align*}
$$

with $M, R, K \in \mathbf{R}^{n \times n}$ being symmetric positive definite, $B_{2} \in \mathbf{R}^{n \times m}$ and $C_{1}, C_{2} \in$ $\mathbf{R}^{l \times n}$ is an instance of (2.1) where $x=(q, M v)$ with $(q, v)$ denoting coordinates on the tangent space $T \mathbf{R}^{n} \cong \mathbf{R}^{n} \times \mathbf{R}^{n}$. The total energy is given by the Hamiltonian

$$
\begin{equation*}
H(q, p)=\frac{1}{2} p^{T} M^{-1} p+\frac{1}{2} q^{T} K q, \quad p=M v . \tag{2.3}
\end{equation*}
$$

Furthermore

$$
J=\left(\begin{array}{cc}
\mathbf{0} & \mathbf{1}_{n \times n}  \tag{2.4}\\
-\mathbf{1}_{n \times n} & \mathbf{0}
\end{array}\right), \quad D=\left(\begin{array}{cc}
\mathbf{0} & \mathbf{0} \\
\mathbf{0} & R
\end{array}\right)
$$

in (2.1) and control and observable matrices in (2.1) and (2.2) are related by

$$
B=\binom{\mathbf{0}}{B_{2}}, \quad C=\left(\begin{array}{cc}
C_{1} K^{-1} & C_{2} \tag{2.5}
\end{array}\right) .
$$

In equation (2.4) and in the following we shall use the notation $\mathbf{1}_{n \times n}$ above to denote the unit matrix of size $n \times n$. Given a solution of (2.2), the energy balance

$$
\frac{d}{d t} H(q(t), M \dot{q}(t))=-2 \gamma(\dot{q}(t))+\dot{q}(t)^{T} B_{1} u(t)
$$

holds where

$$
\gamma(v)=\frac{1}{2} v^{T} R v>0
$$

denotes the Rayleigh dissipation function. As the total energy $H \geq 0$ is bounded from below, it follows that (2.2) with outputs $y=B_{2}^{T} \dot{q}$, i.e., $C=B^{T}$ is a passive state space system with storage function $H$; see [19] for details.
2.1. Balancing transformations. We briefly review the idea of balancing that is due to [21]; see also the textbooks [20,2] and the references therein. As we have pointed out, we shall treat the case of the second-order system (2.2) by considering the equivalent Hamiltonian system (2.1).

For $M, R, K$ in (2.2) being symmetric and positive definite matrices, the firstorder system (2.1) is stable, i.e., all eigenvalues of the constant matrix

$$
A=(J-D) \nabla^{2} H(x), \quad \nabla^{2} H(x)=\left(\begin{array}{cc}
K & \mathbf{0} \\
\mathbf{0} & M^{-1}
\end{array}\right)
$$

are lying in the open left complex half-plane (see Sec. 4). The controllability function

$$
L_{c}(x)=\min _{u \in L^{2}} \int_{-\infty}^{0}|u(t)|^{2} d t, \quad x(-\infty)=0, x(0)=x
$$

then measures the minimum energy that is needed to steer the system from $x(-\infty)=0$ to $x(0)=x$. In turn, the observability function

$$
L_{o}(x)=\int_{0}^{\infty}|y(t)|^{2} d t, \quad x(0)=x, u \equiv 0
$$

measures the control-free energy of the output as the system evolves from $x(0)=x$ to $x(\infty)=0$ (asymptotic stability). It is easy to see that

$$
L_{c}(x)=x^{T} W_{\mathrm{c}}^{-1} x, \quad L_{o}(x)=x^{T} W_{\mathrm{o}} x
$$

where the controllability Gramian $W_{\mathrm{c}}$ and the observability $W_{\mathrm{o}}$ are the unique symmetric solutions of the Lyapunov equations

$$
A W_{\mathrm{c}}+W_{\mathrm{c}} A^{T}=-B B^{T}, \quad A^{T} W_{\mathrm{o}}+W_{\mathrm{o}} A=-Q^{T} Q
$$

with the shorthand $Q=C \nabla^{2} H$. Moore [21] has shown that, if $W_{\mathrm{c}}, W_{\mathrm{o}} \succ 0$ (positive definiteness $=$ complete controllability/observability), there exists a coordinate transformation $x \mapsto T x$ such that the two Gramians become equal and diagonal, ${ }^{1}$

$$
T^{-1} W_{\mathrm{c}} T^{-T}=T^{T} W_{\mathrm{o}} T=\operatorname{diag}\left(\sigma_{1}, \ldots, \sigma_{2 n}\right)
$$

where the Hankel singular values (HSV) $\sigma_{1} \geq \sigma_{2} \geq \ldots \geq \sigma_{2 n}>0$ are independent of the choice of coordinates. We shall assume throughout the paper that our system is

[^1]completely controllable and observable. ${ }^{2}$ A convenient way to express the balancing transformation is due to Moore [21]: Noting that $\Sigma^{2}$ contains the positive eigenvalues of the product $W_{\mathrm{c}} W_{\mathrm{o}}$ we decompose the two Gramians according to
$$
W_{\mathrm{c}}=X X^{T}, \quad W_{\mathrm{o}}=Y Y^{T}
$$
and do a singular value decomposition (SVD) of the matrix $Y^{T} X$, i.e.,
\[

Y^{T} X=U \Sigma V^{T}=\left($$
\begin{array}{cc}
U_{1} & U_{2}
\end{array}
$$\right)\left($$
\begin{array}{cc}
\Sigma_{1} & \mathbf{0}  \tag{2.6}\\
\mathbf{0} & \Sigma_{2}
\end{array}
$$\right)\binom{V_{1}^{T}}{V_{2}^{T}} .
\]

The partitioning $\Sigma_{1}=\operatorname{diag}\left(\sigma_{1}, \ldots, \sigma_{d}\right)$ and $\Sigma_{2}=\operatorname{diag}\left(\sigma_{d+1}, \ldots, \sigma_{2 n}\right)$ indicates which singular values are important and which are negligible. The remaining matrices satisfy $U_{1}^{T} U_{1}=V_{1}^{T} V_{1}=\mathbf{1}_{d \times d}$ and $U_{2}^{T} U_{2}=V_{2}^{T} V_{2}=\mathbf{1}_{r \times r}$ with $r=2 n-d$. In terms of the SVD the balancing transformation $T$ and its inverse $S=T^{-1}$ take the form

$$
\begin{equation*}
T=X V \Sigma^{-1 / 2}, \quad S=\Sigma^{-1 / 2} U^{T} Y^{T} \tag{2.7}
\end{equation*}
$$

It can be readily seen that the balancing transformation leaves the structure of the equations of motion (2.1) unchanged and preserves both stability and passivity. In the balanced variables $z=S x$ our Hamiltonian system reads

$$
\begin{align*}
& \dot{z}(t)=(\tilde{J}-\tilde{D}) \nabla \tilde{H}(z(t))+\tilde{B} u(t) \\
& y(t)=\tilde{C} \nabla \tilde{H}(z(t)) \tag{2.8}
\end{align*}
$$

with the balanced Hamiltonian $\tilde{H}(z)=H(T z)$, i.e.,

$$
\begin{equation*}
\tilde{H}(\xi)=\frac{1}{2} z^{T} \tilde{E} z, \quad \tilde{E}=T^{T} E T \tag{2.9}
\end{equation*}
$$

where $E=\nabla^{2} H(x)$. The transformed coefficients are given by

$$
\begin{equation*}
\tilde{J}=S J S^{T}, \quad \tilde{R}=S R S^{T}, \quad \tilde{B}=S B, \quad \tilde{C}=C S^{T} \tag{2.10}
\end{equation*}
$$

3. Balanced truncation of Hamiltonian systems. Balancing amounts to changing coordinates such that those states that are least influenced by the input also have least influence on the output. Accordingly balanced truncation consists in first balancing the system, and then truncating the least observable and controllable states, which have little effect on the input-output behaviour.
3.1. Strong confinement limit. There are several ways that lead to a truncated (i.e., dimension-reduced) system; standard approaches such as Galerkin (PetrovGalerkin) projection or naive singular perturbation methods, however, fail to preserve the systems inherent Hamiltonian structure. In mechanics, a natural way to restrict a system to a subspace is by means of constraints [23], and, from a physical viewpoint, it makes sense to study the limit of vanishing small Hankel singular values, i.e., to gradually eliminate the least observable and controllable states thereby forcing the system to the limiting controllable and observable subspace.

To make the appearance of the least controllable and observable states in the equations of motion explicit, we scale the HSV uniformly according to

$$
\left(\sigma_{1}, \ldots, \sigma_{d}, \sigma_{d+1}, \ldots, \sigma_{2 n}\right) \mapsto\left(\sigma_{1}, \ldots, \sigma_{d}, \epsilon \sigma_{d+1}, \ldots, \epsilon \sigma_{2 n}\right)
$$

[^2]i.e., in (2.6)-(2.7) we replace $\Sigma_{2}$ by $\epsilon \Sigma_{2}$ and partition the thus obtained balancing matrices accordingly,
\[

S(\epsilon)=\left($$
\begin{array}{cc}
S_{11} & S_{12} \\
\epsilon^{-1 / 2} S_{21} & \epsilon^{-1 / 2} S_{22}
\end{array}
$$\right), \quad T(\epsilon)=\left($$
\begin{array}{cc}
T_{11} & \epsilon^{-1 / 2} T_{12} \\
T_{21} & \epsilon^{-1 / 2} T_{22}
\end{array}
$$\right)
\]

Splitting the state variables $z=\left(z_{1}, z_{2}\right)$ in the same fashion and omitting the free variable in what follows the balanced equations of motion (2.8) take the form

$$
\begin{align*}
\dot{z}_{1}^{\epsilon} & =\left(\tilde{J}_{11}-\tilde{D}_{11}\right) \frac{\partial \tilde{H}^{\epsilon}}{\partial z_{1}}+\frac{1}{\sqrt{\epsilon}}\left(\tilde{J}_{12}-\tilde{D}_{12}\right) \frac{\partial \tilde{H}^{\epsilon}}{\partial z_{2}}+\tilde{B}_{1} u \\
\dot{z}_{2}^{\epsilon} & =\frac{1}{\sqrt{\epsilon}}\left(\tilde{J}_{21}-\tilde{D}_{21}\right) \frac{\partial \tilde{H}^{\epsilon}}{\partial z_{1}}+\frac{1}{\epsilon}\left(\tilde{J}_{22}-\tilde{D}_{22}\right) \frac{\partial \tilde{H}^{\epsilon}}{\partial z_{2}}+\frac{1}{\sqrt{\epsilon}} \tilde{B}_{2} u  \tag{3.1}\\
y^{\epsilon} & =\tilde{C}_{1} \frac{\partial \tilde{H}^{\epsilon}}{\partial z_{1}}+\frac{1}{\sqrt{\epsilon}} \tilde{C}_{2} \frac{\partial \tilde{H}^{\epsilon}}{\partial z_{2}}
\end{align*}
$$

with the scaled Hamiltonian

$$
\tilde{H}^{\epsilon}(z)=\frac{1}{2} z^{T} \tilde{E}^{\epsilon} z, \quad \tilde{E}^{\epsilon}=\left(\begin{array}{cc}
\tilde{E}_{11} & \epsilon^{-1 / 2} \tilde{E}_{12} \\
\epsilon^{-1 / 2} \tilde{E}_{21} & \epsilon^{-1} \tilde{E}_{22}
\end{array}\right) .
$$

Note that the small parameter $\epsilon$ appears singularly in the Hamiltonian $\tilde{H}^{\epsilon}$. If we require the total energy to remain bounded, we conclude that $z_{2}$ goes to zero as $\epsilon \rightarrow 0$ (otherwise the energy would blow up). Hence boundedness of the energy implies that $z_{2}=\mathcal{O}(\sqrt{\epsilon})$ as $\epsilon \rightarrow 0$. It is now helpful to note that $\tilde{H}^{\epsilon}\left(z_{1}, \sqrt{\epsilon} z_{2}\right)$ is uniformly bounded in $\epsilon$, in fact, is independent of $\epsilon$. This suggests to introduce new variables $\left(\zeta_{1}, \zeta_{2}\right)=\left(z_{1}, \epsilon^{-1 / 2} z_{2}\right)$ in terms of which (3.1) becomes

$$
\begin{align*}
\dot{\zeta}_{1}^{\epsilon} & =\left(\tilde{J}_{11}-\tilde{D}_{11}\right) \frac{\partial \tilde{H}}{\partial \zeta_{1}}+\frac{1}{\epsilon}\left(\tilde{J}_{12}-\tilde{D}_{12}\right) \frac{\partial \tilde{H}}{\partial \zeta_{2}}+\tilde{B}_{1} u \\
\dot{\zeta}_{2}^{\epsilon} & =\frac{1}{\epsilon}\left(\tilde{J}_{21}-\tilde{D}_{21}\right) \frac{\partial \tilde{H}}{\partial \zeta_{1}}+\frac{1}{\epsilon^{2}}\left(\tilde{J}_{22}-\tilde{D}_{22}\right) \frac{\partial \tilde{H}}{\partial \zeta_{2}}+\frac{1}{\epsilon} \tilde{B}_{2} u  \tag{3.2}\\
y^{\epsilon} & =\tilde{C}_{1} \frac{\partial \tilde{H}}{\partial \zeta_{1}}+\frac{1}{\epsilon} \tilde{C}_{2} \frac{\partial \tilde{H}}{\partial \zeta_{2}}
\end{align*}
$$

where $\tilde{H}(\zeta)=\tilde{H}^{\epsilon}\left(\zeta_{1}, \sqrt{\epsilon} \zeta_{2}\right)$ is again independent of $\epsilon$.
Limiting equation. Equation (3.2) is an instance of a slow/fast system, and we seek an effective equation for the slow variable $z_{1}=\zeta_{1}$. Let us start with some preliminary considerations: given a solution of (3.2), the following energy balance

$$
\begin{aligned}
\frac{d \tilde{H}}{d t}= & -\left(\frac{\partial \tilde{H}}{\partial \zeta_{1}}\right)^{T} \tilde{D}_{11} \frac{\partial \tilde{H}}{\partial \zeta_{1}}-\frac{2}{\epsilon}\left(\frac{\partial \tilde{H}}{\partial \zeta_{1}}\right)^{T} \tilde{D}_{12} \frac{\partial \tilde{H}}{\partial \zeta_{2}}-\frac{1}{\epsilon^{2}}\left(\frac{\partial \tilde{H}}{\partial \zeta_{2}}\right)^{T} \tilde{D}_{22} \frac{\partial \tilde{H}}{\partial \zeta_{2}} \\
& +\left(\frac{\partial \tilde{H}}{\partial \zeta_{1}}\right)^{T} \tilde{B}_{1} u+\frac{1}{\epsilon}\left(\frac{\partial \tilde{H}}{\partial \zeta_{2}}\right)^{T} \tilde{B}_{2} u
\end{aligned}
$$

holds. Assuming that $\tilde{H}$ remains bounded as $\epsilon$ goes to zero and that the fast dynamics is uniformly hyperbolic for all $\zeta_{1}$, we conclude that (see [24, 25])

$$
\frac{\partial \tilde{H}}{\partial \zeta_{2}} \rightarrow 0 \quad \text { as } \quad \epsilon \rightarrow 0
$$

Latter implies that the dynamics admit a stable invariant manifold given by

$$
\zeta_{2}=-\tilde{E}_{22}^{-1} \tilde{E}_{21} \zeta_{1}
$$

Inserting the last identity into (3.2), we obtain a closed equation for $z_{1}=\zeta_{1}$, viz.,

$$
\begin{align*}
\dot{z}_{1} & =\left(\tilde{J}_{11}-\tilde{D}_{11}\right)\left(\tilde{E}_{11}-\tilde{E}_{12} \tilde{E}_{22}^{-1} \tilde{E}_{21}\right) z_{1}+\tilde{B}_{1} u \\
y & =\tilde{C}_{1}\left(\tilde{E}_{11}-\tilde{E}_{12} \tilde{E}_{22}^{-1} \tilde{E}_{21}\right) z_{1} \tag{3.3}
\end{align*}
$$

Equation (3.3) is Hamiltonian with

$$
\begin{equation*}
\bar{H}\left(z_{1}\right)=\frac{1}{2} z_{1}^{T} \bar{E} z_{1}, \quad \bar{E}=\tilde{E}_{11}-\tilde{E}_{12} \tilde{E}_{22}^{-1} \tilde{E}_{21} \tag{3.4}
\end{equation*}
$$

Notice that $\tilde{J}_{11}=-\tilde{J}_{11}^{T}$ and $\tilde{D}_{11}=\tilde{D}_{11}^{T} \succcurlyeq 0$ are simply the original structure and friction matrices restricted to the subspace of the most controllable and observable states. That is, in the limit of vanishing small HSV the dynamics are confined to the controllable and observable subspace. Note that the confined system is passive if the original system was, i.e., if $C=B^{T}$ in (2.1). Moreover $E=E^{T} \succ 0$ implies $\bar{E}=\bar{E}^{T} \succ 0$ for the Schur complement. We now give a systematic derivation of (3.3); as for the stability issue, we refer to Section 4.

Derivation. The idea of the derivation is to make the transition of (3.2) to the limiting solution (3.3) more lucid by relating solvability conditions of the respective perturbative expansion to the coefficients in the state space system. We suppose that $u \in L^{2}\left(\left[0, \infty\left[, \mathbf{R}^{m}\right)\right.\right.$ and aim at a perturbative expansion of the solutions to (3.2). To this end we observe that the infinitesimal generator $\mathcal{L}^{\epsilon}$ that generates the semigroup of solutions, i.e., the flow of (3.2) can be split acccording to

$$
\mathcal{L}^{\epsilon}=\mathcal{L}_{0}+\frac{1}{\epsilon} \mathcal{L}_{1}+\frac{1}{\epsilon^{2}} \mathcal{L}_{2}
$$

with

$$
\begin{aligned}
\mathcal{L}_{0} & =\left(Z_{11}+\tilde{B}_{1} u\right)^{T} \frac{\partial}{\partial \zeta_{1}} \\
\mathcal{L}_{1} & =Z_{12}^{T} \frac{\partial}{\partial \zeta_{1}}+\left(Z_{21}+\tilde{B}_{2} u\right)^{T} \frac{\partial}{\partial \zeta_{2}} \\
\mathcal{L}_{2} & =Z_{22}^{T} \frac{\partial}{\partial \zeta_{2}}
\end{aligned}
$$

and the shorthand

$$
Z_{i j}=\left(\tilde{J}_{i j}-\tilde{D}_{i j}\right) \frac{\partial \tilde{H}}{\partial \zeta_{j}}
$$

Consider the following Cauchy problem ${ }^{3}$

$$
\begin{equation*}
\partial_{t} v^{\epsilon}(\zeta, t)=\mathcal{L}^{\epsilon} v^{\epsilon}(\zeta, t), \quad v^{\epsilon}(\zeta, 0)=f(\zeta) \tag{3.5}
\end{equation*}
$$

[^3]which is the backward Liouville equation associated with the Hamiltonian system (3.2). The backward equation is fully equivalent to (3.2) by the method of characteristics, i.e., given the solutions $\zeta^{\epsilon}(t)=F_{t}^{\epsilon}(\zeta), F_{0}^{\epsilon}=$ Id of (3.2) with the initial conditions $\zeta^{\epsilon}(0)=\zeta$, the solution to (3.5) is given by $v^{\epsilon}(\zeta, t)=f\left(F_{t}^{\epsilon}(\zeta)\right)$. We seek a perturbative expansion for the Liouville equation that is of the form
$$
v^{\epsilon}=v_{0}+\epsilon v_{1}+\epsilon^{2} v_{2}+\ldots
$$

Plugging the ansatz in the backward equation (3.5) and equating equal powers of $\epsilon$ yields a hierarchy of equations the first three of which are

$$
\begin{align*}
& \mathcal{L}_{2} v_{0}=0  \tag{3.6}\\
& \mathcal{L}_{2} v_{1}=-\mathcal{L}_{1} v_{0}  \tag{3.7}\\
& \mathcal{L}_{2} v_{2}=-\mathcal{L}_{0} v_{0}-\mathcal{L}_{1} v_{1}+\partial_{t} v_{0} . \tag{3.8}
\end{align*}
$$

We proceed step by step: since $\mathcal{L}_{2}$ is a differential operator in $\zeta_{2}$ only and the nullspace of $\tilde{J}_{22}-\tilde{D}_{22}$ is empty, the only functions that solve (3.6) are constant in $\zeta_{2}$, i.e., $v_{0}=v_{0}\left(\zeta_{1}, t\right)$. By the Fredholm alternative [26], equation (3.7) has a solution if and only if the right hand side is orthogonal to the kernel of the $L^{2}$-adjoint $\mathcal{L}_{2}^{*}$ where orthogonality is meant in the $L^{2}$ sense. Since the fast subsystem

$$
\begin{equation*}
\dot{\zeta}_{2}(t)=\left(\tilde{J}_{22}-\tilde{D}_{22}\right)\left(\tilde{E}_{21} \eta+\tilde{E}_{22} \zeta_{2}(t)\right), \quad \zeta_{2}(0)=\zeta_{2}, \tag{3.9}
\end{equation*}
$$

corresponding to $\mathcal{L}_{2}$, is asymptotically stable (see Sec. 4, Cor. 4.5) for each $\zeta_{1}=\eta$ fixed, the dynamics converge exponentially fast to the invariant subspace given by

$$
S=\left\{\zeta_{2} \in \mathbf{R}^{2 n-d} \mid \tilde{E}_{22} \zeta_{2}=-\tilde{E}_{21} \eta\right\}
$$

Solvability of (3.7) therefore requires that the right hand side is zero when we integrate it against any function that is in the nullspace of $\mathcal{L}_{2}^{*}$, i.e., an invariant measure of the fast dynamics. By stability the invariant measure is unique and is given by

$$
d \rho_{\eta}\left(\zeta_{2}\right)=\operatorname{det} \tilde{E}_{22} \delta\left(\tilde{E}_{21} \eta+\tilde{E}_{22} \zeta_{2}\right) d \zeta_{2}
$$

with $\delta(\cdot)$ being the Dirac mass. As $v_{0}$ is independent of $\zeta_{2}$ it follows that

$$
\int_{\mathbf{R}^{2 n-d}} \mathcal{L}_{1} v_{0} d \rho_{\eta}=0
$$

i.e., the solvability condition $\mathcal{L}_{1} v_{0} \perp \operatorname{ker} \mathcal{L}_{2}^{*}$ is met. To solve equation (3.7) we first of all observe that $v_{1}$ must be of the form (see [27])

$$
v_{1}\left(\zeta_{1}, \zeta_{2}, t\right)=\phi\left(\zeta_{1}, \zeta_{2}\right)^{T} \nabla v_{0}\left(\zeta_{1}, t\right)+\psi\left(\zeta_{1}, t\right)
$$

where $\psi \in \operatorname{ker} \mathcal{L}_{2}$ plays no role in what follows, so we set it to zero. Equation (3.7) can now be recast as an equation for $\phi: \mathbf{X} \rightarrow \mathbf{R}^{d}$, the so-called cell problem

$$
\begin{equation*}
\mathcal{L}_{2} \phi=-Z_{12}^{T} \tag{3.10}
\end{equation*}
$$

In (3.5), the initial condition is independent of $\epsilon$, therefore $v_{1}(\zeta, 0)=0$ which leaves the only possible choices $v_{0}=c$ or $\phi=0$. If we exclude the trivial stationary solution $v_{0}$ being constant, consistency of (3.10) requires that $Z_{12}=0$, i.e., the initial conditions
are restricted to the invariant subspace $S$. To conclude, the Fredholm alternative for equation (3.8) entails the solvability condition

$$
\int_{\mathbf{R}^{2 n-d}}\left(\partial_{t} v_{0}-\mathcal{L}_{0} v_{0}-\mathcal{L}_{1} \phi \nabla v_{0}\right) d \rho_{\eta}=0
$$

which, for $\phi=0$, can be recast as

$$
\begin{equation*}
\partial_{t} v_{0}\left(\zeta_{1}, t\right)=\left(\bar{Z}+\tilde{B}_{1} u\right)^{T} \nabla v_{0}\left(\zeta_{1}, t\right) \tag{3.11}
\end{equation*}
$$

with the abbreviation

$$
\bar{Z}=\left(\tilde{J}_{11}-\tilde{D}_{11}\right)\left(\tilde{E}_{11}-\tilde{E}_{12} \tilde{E}_{22}^{-1} \tilde{E}_{21}\right) \zeta_{1}
$$

Upon reinterpetation as a control system, (3.11) equals (3.3).
A note on initial conditions. The derivation of the effective equation (3.11) relies on specific assumptions regarding the initial conditions in (3.5), namely, being independent of $\epsilon$. Exploiting the equivalence of Liouville equation and Hamiltonian system, this implies that the initial conditions $\zeta^{\epsilon}(0)=\zeta$ in the equations of motion (3.2) are independent of $\epsilon$. But this means that $\zeta$ must be restricted to the invariant subspace $S$ since, otherwise, the initial output $y(t=0)$ diverges as $\epsilon \rightarrow 0$ which might produce unphysical behaviour of the system. ${ }^{4}$ If we nevertheless drop the assumption on the initial conditions, the solution of the cell problem (3.10) turns out to be

$$
\phi\left(\zeta_{1}, \zeta_{2}\right)=-\left(\tilde{J}_{12}-\tilde{D}_{12}\right)\left(\tilde{J}_{22}-\tilde{D}_{22}\right)^{-1} \zeta_{2}
$$

where invertibility of $\left(\tilde{J}_{22}-\tilde{D}_{22}\right)$ is due to Corollary 4.5 below, and we have omitted integration constants $C\left(\zeta_{1}\right)$. The Fredholm alternative for (3.8) then yields

$$
\partial_{t} v_{0}=(\bar{W}+\bar{B} u)^{T} \nabla v_{0}
$$

with

$$
\begin{aligned}
\bar{W} & =\left(\tilde{J}_{11}-\tilde{D}_{11}-\left(\tilde{J}_{12}-\tilde{D}_{12}\right)\left(\tilde{J}_{22}-\tilde{D}_{22}\right)^{-1}\left(\tilde{J}_{21}-\tilde{D}_{21}\right)\right) \nabla \bar{H} \\
\bar{B} & =\tilde{B}_{1}-\left(\tilde{J}_{12}-\tilde{D}_{12}\right)\left(\tilde{J}_{22}-\tilde{D}_{22}\right)^{-1} \tilde{B}_{2}
\end{aligned}
$$

That is, for $\phi \neq 0$ the corresponding control system reads

$$
\begin{align*}
\dot{z}_{1} & =(\bar{J}-\bar{D}) \nabla \bar{H}\left(z_{1}\right)+\bar{B} u \\
y & =\bar{C} \nabla \bar{H}\left(z_{1}\right)+\bar{F} u \tag{3.12}
\end{align*}
$$

where

$$
\begin{equation*}
\bar{J}=\frac{1}{2}\left(\bar{L}-\bar{L}^{T}\right) \quad \text { and } \quad \bar{D}=-\frac{1}{2}\left(\bar{L}+\bar{L}^{T}\right) \tag{3.13}
\end{equation*}
$$

are antisymmetric and symmetric parts of the Schur complement of $L=\tilde{J}-\tilde{D}$ and

$$
\begin{align*}
\bar{C} & =\tilde{C}_{1}-\tilde{C}_{2}\left(\tilde{J}_{22}-\tilde{D}_{22}\right)^{-1}\left(\tilde{J}_{21}-\tilde{D}_{21}\right) \\
\bar{F} & =-\tilde{C}_{2}\left(\tilde{J}_{22}-\tilde{D}_{22}\right)^{-1} \tilde{B}_{2} \tag{3.14}
\end{align*}
$$

[^4]are the effective observability coefficients with an additional feed-through term. The extra terms in the effective equation are secular terms that occur if
$$
\frac{\partial \tilde{H}}{\partial \zeta_{2}}=\mathcal{O}(\epsilon)
$$
in (3.2) such that the singular term in the slow equation becomes essentially $\mathcal{O}(1)$. Equation (3.12) is what appears to be the result of a "naive" singular perturbation method that simply recasts the balanced equations (3.1) upon setting $\zeta_{2}=0$. Equations (3.12)-(3.14) have been put forward in [15] based on an energy argument.
3.2. Hard constraints. We shall briefly mention yet another method so as to impose the controllability/observability constraint on the balanced system (2.8) that can be found in [15]. In (2.6), set $\Sigma_{2}=\mathbf{0}$ which renders the equation rank-deficient,
\[

Y^{T} X=\left($$
\begin{array}{ll}
U_{1} & U_{2}
\end{array}
$$\right)\left($$
\begin{array}{cc}
\Sigma_{1} & \mathbf{0} \\
\mathbf{0} & \mathbf{0}
\end{array}
$$\right)\binom{V_{1}^{T}}{V_{2}^{T}}=U_{1} \Sigma_{1} V_{1}^{T}
\]

with $U_{1}^{T} U_{1}=V_{1}^{T} V_{1}=\mathbf{1}_{d \times d}$. We define $T_{1} \in \mathbf{R}^{2 n \times d}$ and $S_{1} \in \mathbf{R}^{d \times 2 n}$ by

$$
T_{1}=X V_{1} \Sigma_{1}^{-1 / 2}, \quad S_{1}=\Sigma_{1}^{-1 / 2} U_{1}^{T} Y^{T}
$$

where the reduced balancing transformation satisfies $S_{1} W_{\mathrm{c}} S_{1}^{T}=\Sigma_{1}=T_{1}^{T} W_{\mathrm{c}} T_{1}$ with $S_{1} T_{1}=\mathbf{1}_{d \times d}$, and we introduce local coordinates $z_{1}=S_{1} x$ on the essential subspace that is defined as the orthogonal complement of the nullspace of $Y^{T} X$. The canonical way to constrain a mechanical system is by restricting the Hamiltonian and the corresponding structure matrix (i.e., the symplectic form). Restricting the Hamiltonian (2.3) according to $x=T_{1} z_{1}$ we thus obtain the restricted system

$$
\begin{align*}
\dot{z}_{1}(t) & =\left(\tilde{J}_{11}-\tilde{D}_{11}\right) \nabla \tilde{H}_{1}\left(z_{1}(t)\right)+\tilde{B}_{1} u(t) \\
y(t) & =\tilde{C}_{1} \nabla \tilde{H}_{1}\left(z_{1}(t)\right) \tag{3.15}
\end{align*}
$$

where $\tilde{J}_{11}=S_{1} J S_{1}^{T}, \tilde{D}_{11}=S_{1} D S_{1}^{T}, \tilde{B}_{1}=S_{1} B, \tilde{C}_{1}=C S_{1}^{T}$ as in (3.3), and

$$
\begin{equation*}
\tilde{H}_{1}\left(z_{1}\right)=\tilde{H}\left(z_{1}, 0\right), \quad \tilde{H}_{1}\left(z_{1}\right)=\frac{1}{2} z_{1}^{T} \tilde{E}_{11} z_{1} \tag{3.16}
\end{equation*}
$$

denotes the restricted Hamiltonian. Note that the thus constrained system shares all properties of (3.3) in terms of structure preservation and passivity.

Remark 3.1. As we will see later on in Section 6 neither (3.12) above nor (3.15) yield controllable approximations of the original dynamics in terms of the associated transfer functions. We mention them just for the sake of completeness.
3.3. On the second-order form of the reduced system. Although the method presented preserves the systems underlying Hamiltonian structure, the same is not necessarily true for the second-order form of the original problem. To see this, we shall write the reduced system as a canonical Hamiltonian system. For this purpose we suppose that $\tilde{J}_{11} \in \mathbf{R}^{d \times d}$ is invertible with $d=2 k$ even, and we note that we can find an invertible transformation $Q \in \mathbf{R}^{d \times d}$ such that [28]

$$
Q \tilde{J}_{11} Q^{T}=\hat{J}, \quad \hat{J}=\left(\begin{array}{cc}
\mathbf{0} & \mathbf{1}_{d \times d} \\
-\mathbf{1}_{d \times d} & \mathbf{0}
\end{array}\right) .
$$

Defining new variables $\xi=Q z_{1}$, the reduced system, say, (3.3) turns into

$$
\begin{aligned}
& \dot{\xi}(t)=(\hat{J}-\hat{D}) \nabla \hat{H}(\xi(t))+\hat{B} u(t) \\
& y(t)=\hat{C} \nabla \hat{H}(\xi(t))
\end{aligned}
$$

with coefficient matrices $\hat{D}=Q \tilde{D}_{11} Q^{T}, \hat{B}=Q \tilde{B}_{1}, \hat{C}=\tilde{C}_{1} Q^{T}$ and the transformed Hamiltonian $\hat{H}(\xi)=\bar{H}\left(Q^{-1} \xi\right)$. Yet, the fact that we can transform the Hamiltonian part to its canonical form does not entail that $\hat{H}$ splits into purely quadratic kinetic and potential energy terms; moreover also friction and input coefficients do not necessarily have the appropriate form (2.4)-(2.5). Therefore the second-order form of the original problem cannot be recovered in general.

One instance in which the second-order structure is actually preserved is when the balancing transformation decays into pure position and momentum parts if we restrict it to the most controllable and observable subspace. This particular case is certainly rather restrictive, however the numerical examples discussed in Section 6 seem to be of this form as the dominant Hankel singular values appear in pairs corresponding to positions and their conjugate momenta.

The idea of balancing positions and momenta separately is exploited in $[4,6,7]$, but it may fail to preserve stability as has been demonstrated in [7].
4. Stability issues. It is well-known that balanced truncation for linear systems preserves asymptotic stability of the original system, and we may ask whether the reduced systems (3.3) and (3.15) preserve stability, too.

We first of all show that the original system is stable: Suppose $J$ and $E$ are real and non-singular $2 n \times 2 n$ matrices where $J=-J^{T}$ is skew-symmetric and $E=E^{T}$ is symmetric. Given a real, positive semi-definite matrix $D \in \mathbf{R}^{2 n \times 2 n}$, we consider a perturbed eigenvalue problem of the form

$$
\begin{equation*}
(J-\mu D) E v=\lambda v \tag{4.1}
\end{equation*}
$$

Theorem 4.1 (Maddocks \& Overton 1995). Suppose that

$$
u^{*} E D E u>0 \quad \forall \text { eigenvectors } u \text { of } J E \text { with pure imaginary eigenvalues. }
$$

Then, for all $\mu>0$ and counting algebraic multiplicity, the number of eigenvalues $\lambda$ of $(J-\mu D) E$ in the open right half-plane $\mathbf{C}^{+}$equals the number of negative eigenvalues of $E$. Furthermore no eigenvalue of $(J-\mu D) E$ is pure imaginary for $\mu>0$.

For the proof using an homotopy argument, we refer to [29]. We have
Lemma 4.2. The system (2.1) with the Hamiltonian (2.3) and the matrices (2.4) is asymptotically stable, i.e., all eigenvalues of $(J-D) E$ lie in the open left half-plane.

Proof. The matrix $E$ is positive definite and symmetric, that is, all its eigenvalues are positive. Then, according to Theorem 4.1, it suffices to prove that $u^{*} E D E u>0$ on the invariant subspace of $J E$ that corresponds to pure imaginary eigenvalues. We prove this by contradiction: Assume the contrary, i.e., there is an eigenvector $u=\left(u_{1}, u_{2}\right)$ of $J E$ with associated imaginary eigenvalue and $u^{*} E D E u=0$. Since

$$
u^{*} E D E u=u_{2}^{*} M^{-1} R M^{-1} u_{2},
$$

we conclude that $u_{2}=0$. But then

$$
J E u=\binom{0}{-K u_{1}} .
$$

Hence $u_{1}=0$, which contradicts that $u$ is an eigenvector.
4.1. Stability preservation. The reasoning in the proof of Lemma 4.2 heavily relies on the specific form of the unbalanced friction matrix $D$, and we cannot use the same trick to prove stability of the dimension-reduced systems (3.3) and (3.15).

In fact we immediately recognize that (asymptotic) stability cannot be preserved, as, e.g., a block diagonal balancing transformation with $k=n$ and $S_{12}=\mathbf{0} \in \mathbf{R}^{n \times n}$ would lead to the reduced system $\dot{z}_{1}=\tilde{B}_{1} u$. Thus the dimension-reduced system is only neutrally stable in general. If, however, $J_{11}$ is non-singular, asymptotic stability is indeed preserved. We shall discuss the two scenarios separately; the case of a singular structure matrix is treated later on in the text.

Nonsingular structure matrix. We suppose that $\tilde{J}_{11}=S_{1} J S_{1}^{T}$ is invertible which implies that the dimension $k$ of the reduced system is even. Let $Q=Q^{T} \succ 0$ denote the Hessian of the reduced Hamiltonian, i.e., either $Q=\tilde{E}_{11}$ the restricted Hessian or $Q=\bar{E}$ the Schur complement. The following can be proved:

Lemma 4.3. Let $\tilde{J}_{11} \in \mathbf{R}^{d \times d}$ be invertible and $\tilde{D}_{11}$ be given as below (3.15). For any $d \times d$ matrix $Q=Q^{T} \succ 0$ the matrix $P=\left(\tilde{J}_{11}-\tilde{D}_{11}\right) Q$ is stable, i.e., all eigenvalues of $P$ lie in the open left half-plane.

Proof. Since $Q$ is symmetric, positive definite, we may write

$$
Q^{1 / 2}\left(\tilde{J}_{11} Q\right) Q^{-1 / 2}=Q^{1 / 2} \tilde{J}_{11} Q^{1 / 2}
$$

i.e., $\tilde{J}_{11} Q$ is similar to a skew-symmetric matrix and so its spectrum lies entirely on the imaginary axis. Moreover $\tilde{J}_{11}$ is non-singular, so we can exclude 0 to be an eigenvalue. Let $v$ be an eigenvector of $\tilde{J}_{11} Q$ with pure imaginary eigenvalue. The vector

$$
w=Q v
$$

solves the eigenvalue problem

$$
Q^{1 / 2} \tilde{J}_{11} Q^{1 / 2} w=\lambda w, \quad \lambda \neq 0
$$

Upon multiplying the last equation with $w^{*}$ from the left, we obtain

$$
w^{*} Q^{1 / 2} \tilde{J}_{11} Q^{1 / 2} w=\lambda\|w\|^{2}
$$

Since the right hand side of this equation is non-zero by construction, we have

$$
\begin{equation*}
w^{*} Q^{1 / 2} \tilde{J}_{11} Q^{1 / 2} w \neq 0 \tag{4.2}
\end{equation*}
$$

Theorem 4.1 guarantees that $P=\left(\tilde{J}_{11}-\tilde{D}_{11}\right) Q$ is stable if $v^{*} Q \tilde{D}_{11} Q v>0$ where $v$ is an arbitrary eigenvector of $\tilde{J}_{11} Q$. We assume that $v^{*} Q \tilde{D}_{11} Q v=0$ and aim for a contradiction. The corresponding condition for $w$ reads

$$
w^{*} Q^{1 / 2} \tilde{D}_{11} Q^{1 / 2} w=0
$$

Since $\tilde{D}_{11}=S_{12} R S_{12}^{T}$ with $R=R^{T} \succ 0$ we conclude that $Q^{1 / 2} w$ lies in the nullspace of $S_{12}^{T}$. But $\tilde{J}_{11}=S_{11} S_{12}^{T}-S_{12} S_{11}^{T}$, and therefore the left hand side of (4.2) is zero which yields the contradiction. $\square$

Corollar 4.4. The reduced Hamiltonian systems (3.3) and (3.15) are stable whenever $\tilde{J}_{11}$ is non-singular.

Singular structure matrix. The requirement that the restricted structure matrix is invertible is rather restrictive. For instance, $k$ may be odd in case of which $\tilde{J}_{11} \in \mathbf{R}^{k \times k}$ has at least one eigenvalue zero. As we have argued it may even happen that $\tilde{J}_{11}-\tilde{D}_{11}$ vanishes identically, although $S_{1}$ has maximum rank. To treat the general case general we assume that the $d \times d$ matrix

$$
\tilde{J}_{11}-\tilde{D}_{11}=S_{1}(J-D) S_{1}^{T}
$$

has reduced rank $d-s$. That is, the autonomous system (i.e., for $u=0$ ) admits certain conserved quantities $C_{i}: \mathbf{R}^{d} \rightarrow \mathbf{R}, i=1, \ldots, s$ (Casimirs) that satisfy

$$
\nabla C_{i}\left(z_{1}\right)^{T}\left(\tilde{J}_{11}-\tilde{D}_{11}\right)=0, \quad i=1, \ldots, s
$$

Clearly, the $C_{i}$ are of the form $C_{i}\left(z_{1}\right)=b_{i}^{T} z_{1}$. Accordingly the dynamics are unstable (i.e., not asymptotically stable) in the directions orthogonal to the level sets of the $C_{i}\left(z_{1}\right)$. The system is still completely controllable though, and it is easy to see that we can transform the equations of motion to assume the following form [16]

$$
\begin{aligned}
\dot{\xi} & =(\hat{J}(C)-\hat{D}(C)) \frac{\partial \hat{H}}{\partial \xi}+\hat{B}_{1}(C) u \\
\dot{C} & =\hat{B}_{2}(C) u
\end{aligned}
$$

where $\xi=\left(\xi_{1}, \ldots, \xi_{d-s}\right)$ are local coordinates on the hyperplane $C\left(z_{1}\right)=C$, and $\hat{J}-\hat{D}$ is the restriction of $\tilde{J}_{11}-\tilde{D}_{11}$ to the hyperplane. Since $\hat{J}$ is invertible and $\hat{D} \succcurlyeq 0$, the restricted system remains stable along the $\xi$-direction.
4.2. Invariant manifold. Recalling the singular perturbation argument from Section 3.1, it remains to prove that the subspace of uncontrollable/unobservable states to which the dynamics collapses as the small HSV go to zero is indeed a stable invariant manifold. In other words, we have to show that the eigenvalues of the matrix $\left.\left(\tilde{J}_{22}-\tilde{D}_{22}\right) \tilde{E}_{22}\right)$ have strictly negative real part. We suppose the original system (2.1) is minimal such that all the HSV in (2.6) are strictly positive. In accordance with (2.7) we define $T_{2} \in \mathbf{R}^{2 n \times(2 n-d)}$ and $S_{2} \in \mathbf{R}^{(2 n-d) \times 2 n}$ by

$$
T_{2}=X V_{2} \Sigma_{2}^{-1 / 2}, \quad S_{2}=\Sigma_{2}^{-1 / 2} U_{2}^{T} Y^{T}
$$

which are the balancing transformation on the fast (i.e., least controllable/observable) subspace. By positivity of $E=\nabla^{2} H(x)$ the matrix $\tilde{E}_{22}=T_{2}^{T} E T_{2}$ is symmetric positive definite. Employing the notation $S_{2}=\left(S_{21}, S_{22}\right)$ while assuming that $S_{22} \neq \mathbf{0}$ structure and friction matrix on the fast subspace take the form

$$
\tilde{D}_{22}=S_{22} R S_{22}^{T}, \quad \tilde{J}_{22}=S_{21} S_{22}^{T}-S_{22} S_{21}^{T}
$$

The following is a straight consequence of Lemma 4.3.
Corollar 4.5. Let $\tilde{J}_{22}$ be non-singular and $\tilde{D}_{22} \succcurlyeq 0$. Then the fast subsystem (3.9) admits a unique stable invariant manifold $S$.

Proof. Obviously $S$ is an invariant manifold of (3.9). To show that it is unique and asymptotically stable it suffices to show that both $\tilde{J}_{22}-\tilde{D}_{22}$ and $\left(\tilde{J}_{22}-\tilde{D}_{22}\right) \tilde{E}_{22}$ are stable and have no have no eigenvalues on the imaginary axis. Adapting the proof of Lemma 4.3 with $Q=\mathbf{1}$ or $Q=\tilde{E}_{22}$, respectively, yields the result. $\square$

The last statement guarantees that (3.2) is hyperbolic in the sense that, as $\epsilon$ goes to zero, the fast dynamics contracts exponentially to its stationary point conditional
on the fixed value of the slow variable. By invertibility of $\tilde{J}_{22}-\tilde{R}_{22}$, the family of the stationary points is uniquely determined by $S$.

REMARK 4.6. In the stability proofs above we have taken advantage of the fact that the friction matrix $R \in \mathbf{R}^{n \times n}$ is symmetric and positive definite. This is certainly more than what Theorem 4.1 demands since stability only requires that the friction acts on the eigenspaces of the pure imaginary eigenvalues of $\tilde{J}_{11} Q$ or $\tilde{J}_{22} Q$. For future research it would be desirable to generalize the proofs to the case of $R=R^{T}$ being lowrank. Yet another extension which, however, is not at all covered by Theorem 4.1 would involve dissipation coming from gyroscopic forces in which case the friction matrix would be skew-symmetric thereby leading to Hamiltonian eigenvalue problems.

All other results regarding the singular perturbation argument easily carry over to these scenarios provided the system is stable.
5. Minimal realization. Before we conclude with some numerical examples we demonstrate that the dimension-reduced system (3.3) provides a minimal realization of the original system when the weakly controllable/observable modes become completely uncontrollable/unobservable. This is not completely obvious as the coefficients of the Hamiltonian system are different from those of a standard balanced and truncated system (see Sec. 6) that is known to amount to the minimal realization if the respective Hankel singular values are exactly zero. What we prove here is that, in the limit of vanishing small Hankel singular values, the transfer function of the original Hamiltonian system (2.1) converges to the reduced transfer function associated with (3.3). The transfer function associated with (2.1) reads

$$
\begin{equation*}
G(s)=C E(s-(J-D) E)^{-1} B, \quad E=\nabla^{2} H \tag{5.1}
\end{equation*}
$$

It can be regarded as the solution to (2.1) in the Laplace domain with zero initial conditions considered as a input/output mapping $G: L^{2}\left(\left[0, \infty\left[, \mathbf{R}^{m}\right) \rightarrow L^{2}\left(\left[0, \infty\left[, \mathbf{R}^{l}\right)\right.\right.\right.\right.$; see [1] for details. Provided that $(J-D) E$ is stable, the transfer function is analytic in the open right half-plane, and the $H^{\infty}$ norm of $G$ is defined as the supremum of the largest singular value of the transfer function on the imaginary axis [2],

$$
\|G\|_{\infty}=\sup _{\omega \in \mathbf{R}}\left\{\sigma_{\max }(G(i \omega))\right\}
$$

Convergence. We now prove convergence of (2.1) to the reduced system (3.3) in the $H^{\infty}$ norm in terms of the corresponding transfer functions. The transfer function of the singularly perturbed system (3.2) given by

$$
\begin{equation*}
G^{\epsilon}(s)=\tilde{C}^{\epsilon} \tilde{E}\left(s-\left(\tilde{J}^{\epsilon}-\tilde{D}^{\epsilon}\right) \tilde{E}\right)^{-1} \tilde{B}^{\epsilon} \tag{5.2}
\end{equation*}
$$

where the scaled coefficient matrices read

$$
\tilde{J}^{\epsilon}-\tilde{D}^{\epsilon}=\left(\begin{array}{cc}
\tilde{J}_{11}-\tilde{D}_{11} & \epsilon^{-1}\left(\tilde{J}_{12}-\tilde{D}_{12}\right) \\
\epsilon^{-1}\left(\tilde{J}_{21}-\tilde{D}_{21}\right) & \epsilon^{-2}\left(\tilde{J}_{22}-\tilde{D}_{22}\right)
\end{array}\right)
$$

and

$$
\tilde{B}^{\epsilon}=\binom{\tilde{B}_{1}}{\epsilon^{-1} \tilde{B}_{2}}, \quad \tilde{C}^{\epsilon}=\left(\begin{array}{cc}
\tilde{C}_{1} & \epsilon^{-1} \tilde{C}_{2}
\end{array}\right)
$$

Note that $\tilde{E}=\nabla^{2} \tilde{H}$ does not depend on $\epsilon$. Moreover, $G^{\epsilon=1}$ coincides with the transfer function $G$ of the original system (2.1). We have

Theorem 5.1. Let $G^{\epsilon}$ be the scaled transfer function (5.2), and denote by $\bar{G}$ the reduced transfer function of the limit system (3.3), i.e.,

$$
\begin{equation*}
\bar{G}(s)=\tilde{C}_{1} \bar{E}\left(s-\left(\tilde{J}_{11}-\tilde{D}_{11}\right) \bar{E}\right)^{-1} \tilde{B}_{1} \tag{5.3}
\end{equation*}
$$

with $\bar{E}=\tilde{E}_{11}-\tilde{E}_{12} \tilde{E}_{22}^{-1} \tilde{E}_{21}$ being the Schur complement of $\tilde{E}$. Then

$$
\left\|G^{\epsilon}-\bar{G}\right\|_{\infty} \rightarrow 0 \quad \text { as } \quad \epsilon \rightarrow 0
$$

Proof. We show that $G^{\epsilon}=\bar{G}+\mathcal{O}(\epsilon)$ by Taylor expanding $G^{\epsilon}$ around $\epsilon=0$; the calculation is tedious but straightforward. First of all recall that the inverse of a partitioned matrix $X \in \mathbf{R}^{2 n \times 2 n}$ can be written as

$$
\left(\begin{array}{ll}
X_{11} & X_{12} \\
X_{21} & X_{22}
\end{array}\right)^{-1}=\left(\begin{array}{cc}
Y^{-1} & -X_{11}^{-1} X_{12} Z^{-1} \\
-X_{22}^{-1} X_{21} Z^{-1} & Z^{-1}
\end{array}\right)
$$

with the shorthands $Y=X_{11}-X_{12} X_{22}^{-1} X_{21}$ and $Z=X_{22}-X_{21} X_{11}^{-1} X_{12}$. Hence

$$
\left(\begin{array}{cc}
s-\tilde{A}_{11}^{\epsilon} & \tilde{A}_{12}^{\epsilon} \\
\tilde{A}_{21}^{\epsilon} & s-\tilde{A}_{22}^{\epsilon}
\end{array}\right)^{-1}=:\left(\begin{array}{ll}
W_{11}^{\epsilon} & W_{12}^{\epsilon} \\
W_{21}^{\epsilon} & W_{22}^{\epsilon}
\end{array}\right)
$$

with $\tilde{A}^{\epsilon}=\left(\tilde{J}^{\epsilon}-\tilde{D}^{\epsilon}\right) \tilde{E}$. This yields

$$
\begin{align*}
G^{\epsilon}(s) & =\underbrace{\left(\tilde{C}_{1} \tilde{E}_{11}+\frac{1}{\epsilon} \tilde{C}_{2} \tilde{E}_{21}\right) W_{11}^{\epsilon} \tilde{B}_{1}}_{(i)}+\underbrace{\left(\tilde{C}_{1} \tilde{E}_{12}+\frac{1}{\epsilon} \tilde{C}_{2} \tilde{E}_{22}\right) W_{21}^{\epsilon} \tilde{B}_{1}}_{(i i i)} \\
& +\underbrace{\frac{1}{\epsilon}\left(\tilde{C}_{1} \tilde{E}_{11}+\frac{1}{\epsilon} \tilde{C}_{2} \tilde{E}_{21}\right) W_{12}^{\epsilon} \tilde{B}_{2}}_{(i v)}+\underbrace{}_{\left(i i{ }^{\frac{1}{\epsilon}\left(\tilde{C}_{1} \tilde{E}_{12}+\frac{1}{\epsilon} \tilde{C}_{2} \tilde{E}_{22}\right) W_{22}^{\epsilon} \tilde{B}_{2}}\right.} \tag{5.4}
\end{align*}
$$

where, e.g., $W_{11}^{\epsilon}$ is given by

$$
\begin{aligned}
W_{11}^{\epsilon} & =\left[s-\left(\tilde{J}_{11}-\tilde{D}_{11}\right) \tilde{E}_{11}-\frac{1}{\epsilon}\left(\tilde{J}_{12}-\tilde{D}_{12}\right) \tilde{E}_{21}\right. \\
& -\left(\left(\tilde{J}_{11}-\tilde{D}_{11}\right) \tilde{E}_{12}+\frac{1}{\epsilon}\left(\tilde{J}_{12}-\tilde{D}_{12}\right) \tilde{E}_{22}\right) \\
& \times\left(s-\frac{1}{\epsilon}\left(\tilde{J}_{21}-\tilde{D}_{21}\right) \tilde{E}_{12}-\frac{1}{\epsilon^{2}}\left(\tilde{J}_{22}-\tilde{D}_{22}\right) \tilde{E}_{22}\right)^{-1} \\
& \left.\times\left(\frac{1}{\epsilon}\left(\tilde{J}_{21}-\tilde{D}_{21}\right) \tilde{E}_{11}+\frac{1}{\epsilon^{2}}\left(\tilde{J}_{22}-\tilde{D}_{22}\right) \tilde{E}_{21}\right)\right]^{-1}
\end{aligned}
$$

The result follows upon Taylor expansion around $\epsilon=0$; we give only a short sketch for the derivation: the first term in (5.4) yields up to terms of order $\epsilon$

$$
(i) \approx \tilde{C}_{1} \tilde{E}_{11}\left[s-\left(\tilde{J}_{11}-\tilde{D}_{11}\right)\left(\tilde{E}_{11}-\tilde{E}_{12} \tilde{E}_{22}^{-1} \tilde{E}_{21}\right)\right]^{-1} \tilde{B}_{1}
$$

Expanding the second term, we find up to order $\epsilon$ :

$$
(i i) \approx-\tilde{C}_{1} \tilde{E}_{12} \tilde{E}_{22}^{-1} \tilde{E}_{21}\left[s-\left(\tilde{J}_{11}-\tilde{D}_{11}\right)\left(\tilde{E}_{11}-\tilde{E}_{12} \tilde{E}_{22}^{-1} \tilde{E}_{21}\right)\right]^{-1} \tilde{B}_{1}
$$

All remaining terms in $(i)-(i i)$ and the expansion of $(i i i)-(i v)$ are formally of order $\epsilon$. Thus the transfer function of the full system can be recast as

$$
G^{\epsilon}(s)=\tilde{C}_{1} \bar{E}\left(s-\left(\tilde{J}_{11}-\tilde{E}_{11}\right) \bar{E}\right)^{-1} \tilde{B}_{1}+\epsilon \rho^{\epsilon}
$$

where $\rho^{\epsilon}$ is uniformly bounded in $\epsilon$, since all matrices involved remain non-singular. Hence we can bound $\rho^{\epsilon}$ by its largest singular value (that is bounded in $\epsilon$ ). Using the triangle inequality we therefore conclude that $\left\|G^{\epsilon}-\bar{G}\right\|_{\infty} \rightarrow 0$ as $\epsilon \rightarrow 0$.

Error bounds. The above result shows that the $H^{\infty}$ error of the reduced system is of order $\epsilon$. In fact it follows from standard perturbation analysis [12] for the transfer functions that the reduced system (3.3) with zero initial condition satisfies

$$
\begin{equation*}
\left\|G^{\epsilon}-\bar{G}\right\|_{\infty}<2 \epsilon\left(\sigma_{d+1}+\ldots+\sigma_{2 n}\right) \tag{5.5}
\end{equation*}
$$

which is the usual upper balanced truncation bound that is due to Glover [22] and that holds for both standard balanced truncation and singular perturbation approaches. Following [12], latter entails also the strong confinement method of Section 3.1. However, the bound does not hold for the system including secular terms, (3.12), or the "naively" constrained system (3.15) as the numerical results below demonstrate.

In principle it is even possible to proceed in the expansion so as to obtain a sharper bound on the error in terms of the remainder $\rho^{\epsilon}$ for $\epsilon=1$. Computing these terms, however, is tedious and the resulting expressions involve a lot of matrix operations (especially inverting large matrices) that are also difficult to compute numerically.

Remark 5.2. By construction of the reduced system as the strong confinement limit of the original one in the time domain, the reduced system (3.3) gives an approximation of the flow of (2.1) for all initial conditions that lie on (or sufficiently close to) the system's invariant subspace; in particular, zero is an admissible initial condition. On the other hand the tacit assumption, when it comes to the transfer function (using Laplace or Fourier transforms), is that the initial condition is always zero. Consequently, convergence on the level of the equations of motion for nonzero initial conditions implies convergence of the transfer function, Theorem 5.1, but the converse is not true. Yet we believe that Theorem 5.1 deserves special attention and is a result in its its own right as its proof reveals that the error of the reduced system is of the order of the largest small Hankel singular value.
6. Numerical illustration. For second-order systems of the form

$$
\begin{aligned}
M \ddot{q}(t)+R \dot{q}(t)+K q(t) & =B_{2} u(t) \\
y(t) & =C_{1} q(t)+C_{2} \dot{q}(t)
\end{aligned}
$$

with $M, R, K \in \mathbf{R}^{n \times n}$ being all symmetric positive definite, $B_{2} \in \mathbf{R}^{n \times m}$ and $C_{i} \in$ $\mathbf{R}^{l \times n}$, we compare four different model reduction methods: standard balanced truncation, reduction by strong confinement (Sec. 3.1), the method of [15] containing the secular terms (end of Sec. 3.1) and "hard" constraints (Sec. 3.2). For this purpose we recast the second-order system as the equivalent Hamiltonian system

$$
\begin{aligned}
\dot{x}(t) & =(J-D) \nabla H(x(t))+B u(t) \\
y(t) & =C \nabla H(x(t)),
\end{aligned}
$$

with $x=(q, M \dot{q})$ and the Hamiltonian function

$$
H=\frac{1}{2} x_{2}^{T} M^{-1} x_{2}+\frac{1}{2} x_{1}^{T} K x_{1} .
$$



Fig. 6.1. Hankel singular values (left panel) and spectral norms of 8-th order approximations (right panel) of the 48-dimensional building model. The dashed line shows the error bound.

Choosing $J=-J^{T}$ to be the canonical structure matrix

$$
J=\left(\begin{array}{cc}
\mathbf{0} & \mathbf{1}_{n \times n} \\
-\mathbf{1}_{n \times n} & \mathbf{0}
\end{array}\right) \in \mathbf{R}^{2 n \times 2 n}
$$

the definition of the remaining coefficients follows accordingly. For benchmarking we choose two model systems from structural mechanics: the international space station (ISS) model and the building model that are both taken from [30]. The comparison is made by computing the spectral norm of $\delta \bar{G}_{d}=G-\bar{G}_{d}$ in the frequency domain where $\bar{G}_{d}: L^{2}\left(\left[0, \infty\left[, \mathbf{R}^{m}\right) \rightarrow L^{2}\left(\left[0, \infty\left[, \mathbf{R}^{l}\right)\right.\right.\right.\right.$ stands for the $d$-th order transfer function associated with one of the reduction schemes. For instance, for standard balanced truncation the reduced transfer function reads

$$
\bar{G}_{d}(s)=\tilde{F}_{1}\left(s-\tilde{A}_{11}\right)^{-1} \tilde{B}_{1}
$$

where $\tilde{F}_{1} \in \mathbf{R}^{l \times d}$ consists of the first $d$ columns of $F=C \nabla^{2} H$ after balancing, $\tilde{A}_{11} \in \mathbf{R}^{d \times d}$ denotes the upper left $d \times d$ block of $A=(J-D) \nabla^{2} H$ after balancing and $\tilde{B}_{1} \in \mathbf{R}^{d \times k}$ is simply the balanced and truncated input matrix $B$; the transfer function of the full system is given by (5.1).

We start by our comparison by computing the Hankel singular values and spectral norms for the building model $(m=l=1$ and $n=24)$. Figure 6.1 shows the error of 8 -th order approximations using standard balanced truncation (Galerkin or PetrovGalerkin projection) and the confined system (3.3). As expected, both methods meet the usual upper $H^{\infty}$ bound (5.5). The reduced model involving the secular terms and the constrained system yield similar results for low-order approximations (slightly worse though), but it turns out that the error exceeds the upper bound as the order of the reduced model is increased (see Fig. 6.2); compare also the numerical studies in the recent article [7]. As for the secular system a possible explanation for this behaviour are inconsistent initial conditions, for the definition of the transfer function relies on zero initial conditions; cf. the discussion of the cell problem at the end of Section 3.1. The same effect is observed for the 270-dimensional ISS model ( $m=$ $l=3$ and $n=135$ ): increasing the number of modes in the approximant renders the spectral norm of the error to exceed the balanced truncation bound as Figure 6.4 indicates. The corresponding Hankel singular values together with the Galerkinprojected and the confined approximants for $d=18$ are depicted in Figure 6.3. It is


FIG. 6.2. Building model: comparison of balanced truncation / confinement (left panel) and secular system / constraint (right panel) for an approximant of order $d=18$.


Fig. 6.3. First 50 HSV (left panel) and spectral norms of 18 -th order approximations (right panel) of the 270-dimensional ISS model. The dashed line shows the balanced truncation bound.
interesting to note that the confined system yields a slightly better approximation in the low-frequency regime than does standard balanced truncation; see also Figures 6.1 and 6.2 in which the effect is even more pronounced. This behaviour can be easily explained by referring to the (time-domain) perturbation approach of Section 3.1 that was about approximating the slowest motion of a slow/fast system by eliminating the fast vibrational modes.

Acknowledgment. We thank Volker Mehrmann, Tatjana Stykel and Timo Reis for valuable discussions and comments. We benefited greatly from the remarks of the anonymous referees that helped to improve the first version of this article and we are grateful for pointing out reference [8]. The work of CH is supported by the DFG Research Center Matheon "Mathematics for Key Technologies" (FZT86) in Berlin. VMV holds a scholarship from the Berlin Mathematical School (BMS).

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Fig. 6.4. ISS model: comparison of balanced truncation / confinement (left panel) and secular system / constraint (right panel) for an approximant of order $d=26$.
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[^1]:    ${ }^{1} T$ is a so-called a contragredient transformation; see [22] for details.

[^2]:    ${ }^{2}$ If the system is not minimal, i.e., completely controllable and observable, a minimal realization has to be computed in advance by Kalman decomposition [1].

[^3]:    ${ }^{3}$ We assume that $\mathcal{L}^{\epsilon}$ is equipped with appropriate boundary conditions. More precisely, we consider (3.5) on all $\mathbf{R}^{2 n}$ and require that $\tilde{H}(\zeta)$ grows "sufficiently" fast as $|\zeta| \rightarrow \infty$; assuming $\tilde{H}$ to be strictly convex is certainly sufficient, but we have to exclude $\nabla^{2} \tilde{H}$ being only semidefinite.

[^4]:    ${ }^{4}$ Alternatively, we can scale $\zeta=\zeta^{\epsilon}$ suitably such that $\zeta^{\epsilon} \rightarrow\left(\zeta_{1},-\tilde{E}_{22}^{-1} \tilde{E}_{21} \zeta_{1}\right)$ as $\epsilon \rightarrow 0$.

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